

Investigation of meteor models with numerical simulation

Toying with models of meteor flight for a deeper understanding

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Overview

Motivation

All work and no play makes Jack a dull boy...

English proverb

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 - ▶ estimate meteor flux
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- ▶ with it we can
 - ▶ estimate meteor flux
 - ▶ validate physical models
 - ▶ investigate r and s indices
 - ▶ ...
- ▶ let's go beyond statistics
- ▶ look at variations in meteors' properties

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 - ▶ ...
 - ▶ **anything**
4. look at the dataset

Simulation

Simulation of flight

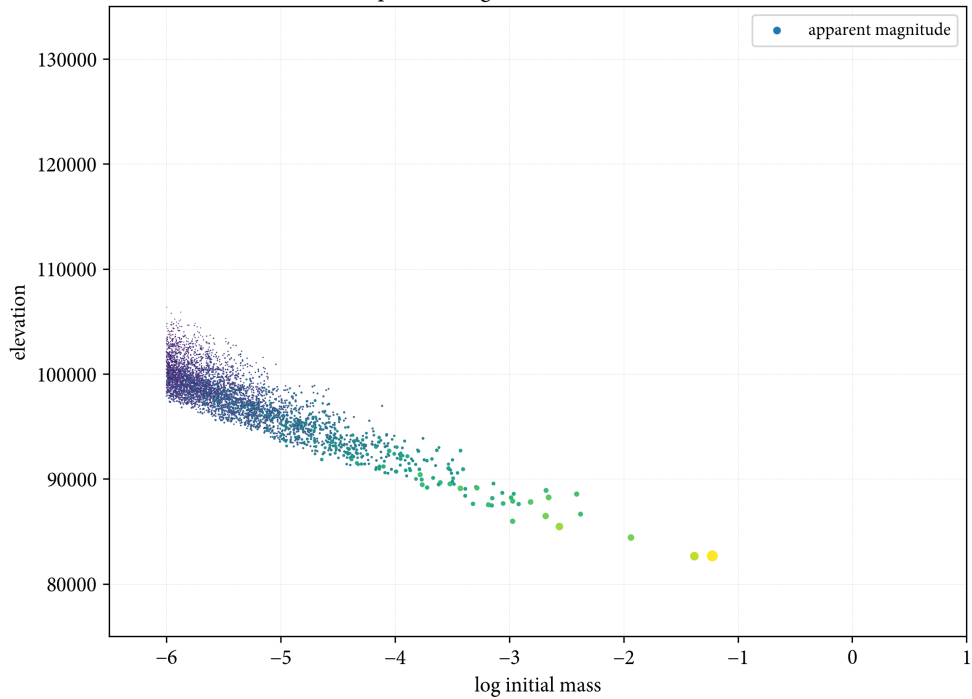
Standard model – **Öpik** (1955)

- ▶ single spherical body
- ▶ continuous thermal ablation
- ▶ no ionization or trail emission

3D simulation on spherical Earth

- ▶ Runge–Kutta integrator
- ▶ **snapshots** taken N times per second
- ▶ until complete ablation (or...)

Tepličné – $\log \text{ initial mass} \times \text{elevation}$



Prerequisites

Grid generator

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- ▶ simulate, observe, analyze...
- ▶ look for changes in the output

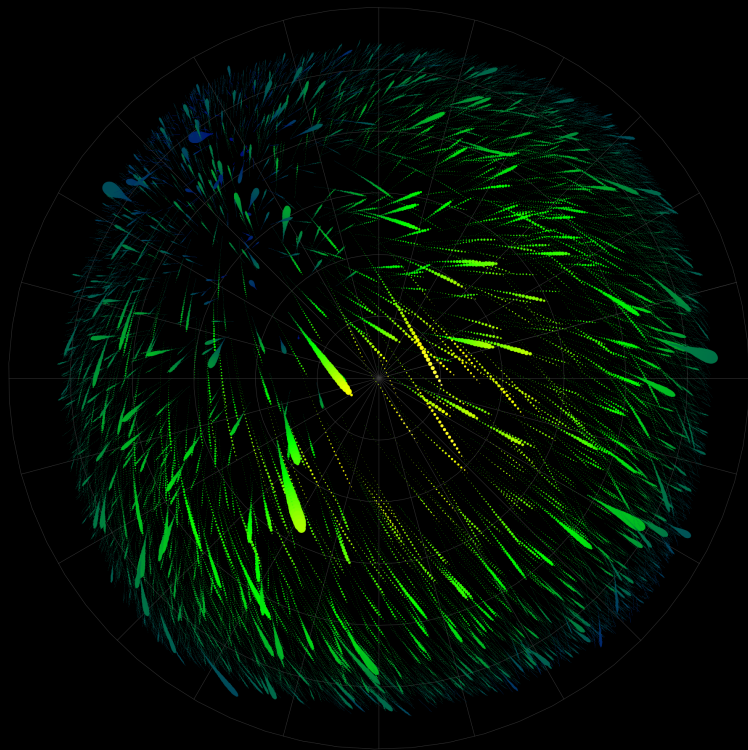
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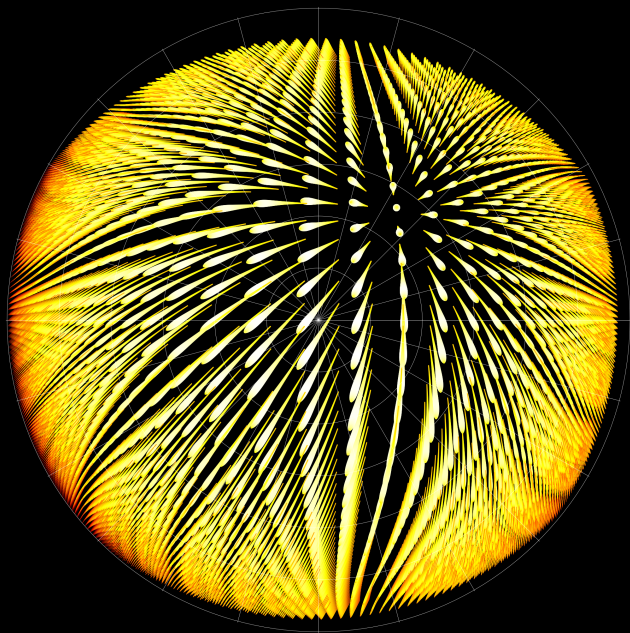
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- ▶ ...except one...
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We may look at the **brightest frame** or **entire trail**





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- ▶ scatter plot
- ▶ colour by varied property
- ▶ emphasize **max-light frame**

What is it good for?

- ▶ predict properties of **any** incoming particle
- ▶ improve ablation and flight models
- ▶ try to find something interesting in the data

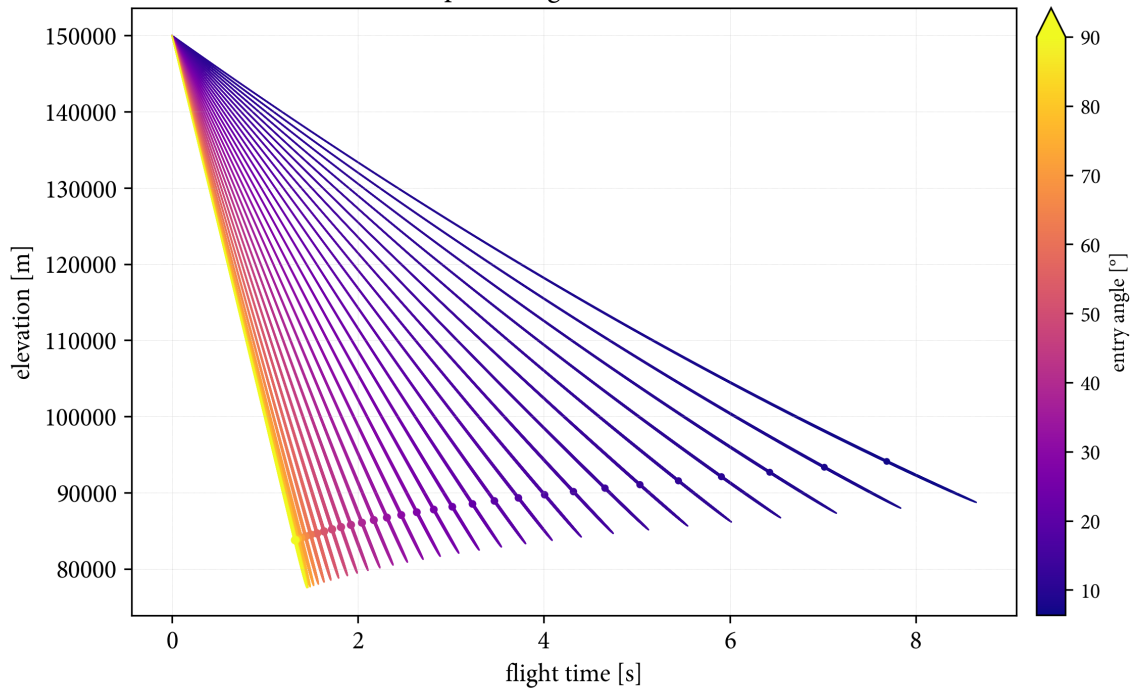
Entry angle

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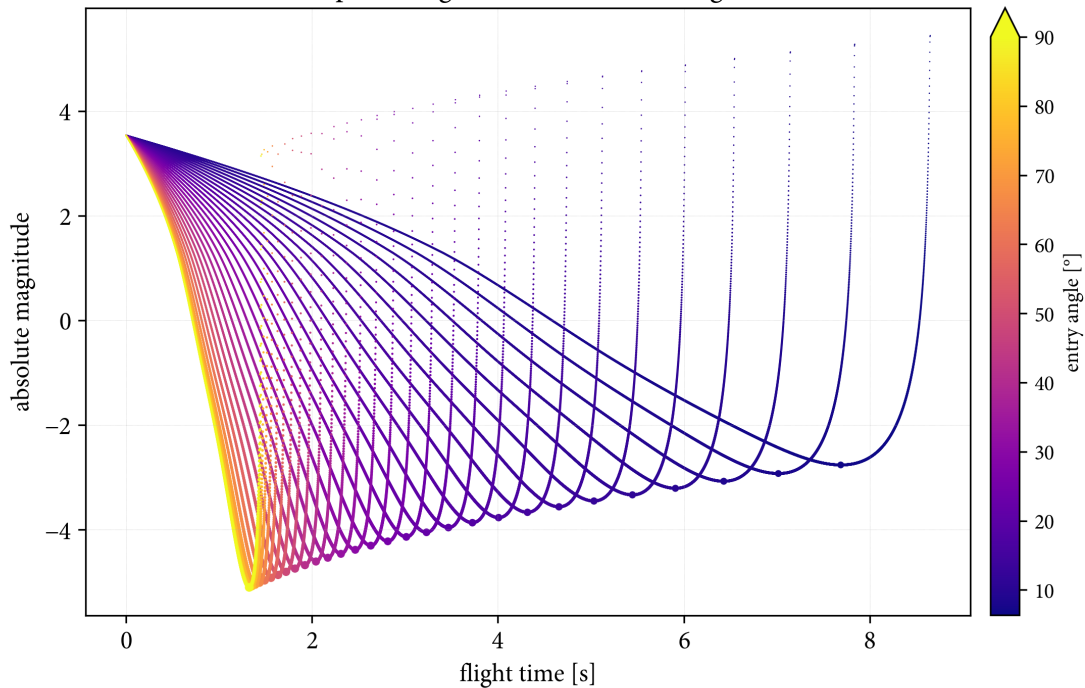
varying **entry angle**

- ▶ Perseids
 - ▶ material, velocity, ...
- ▶ mass 1 g
- ▶ θ between 10° and 90°

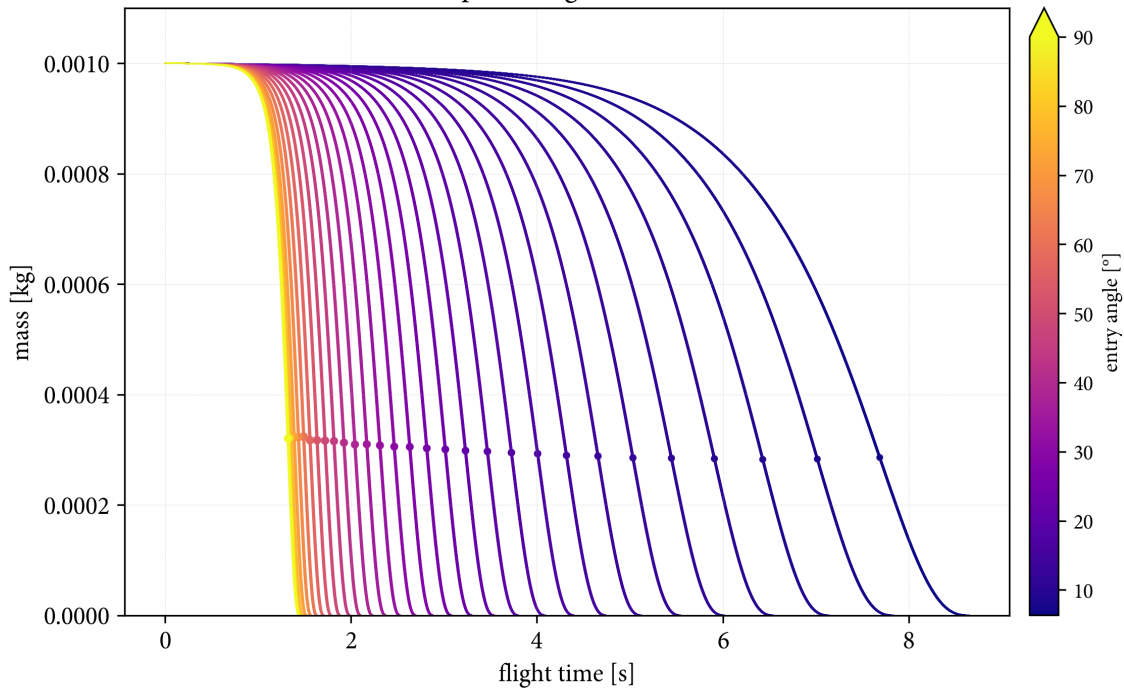
North pole – flight time \times elevation



North pole – flight time \times absolute magnitude



North pole – flight time \times mass



Mass

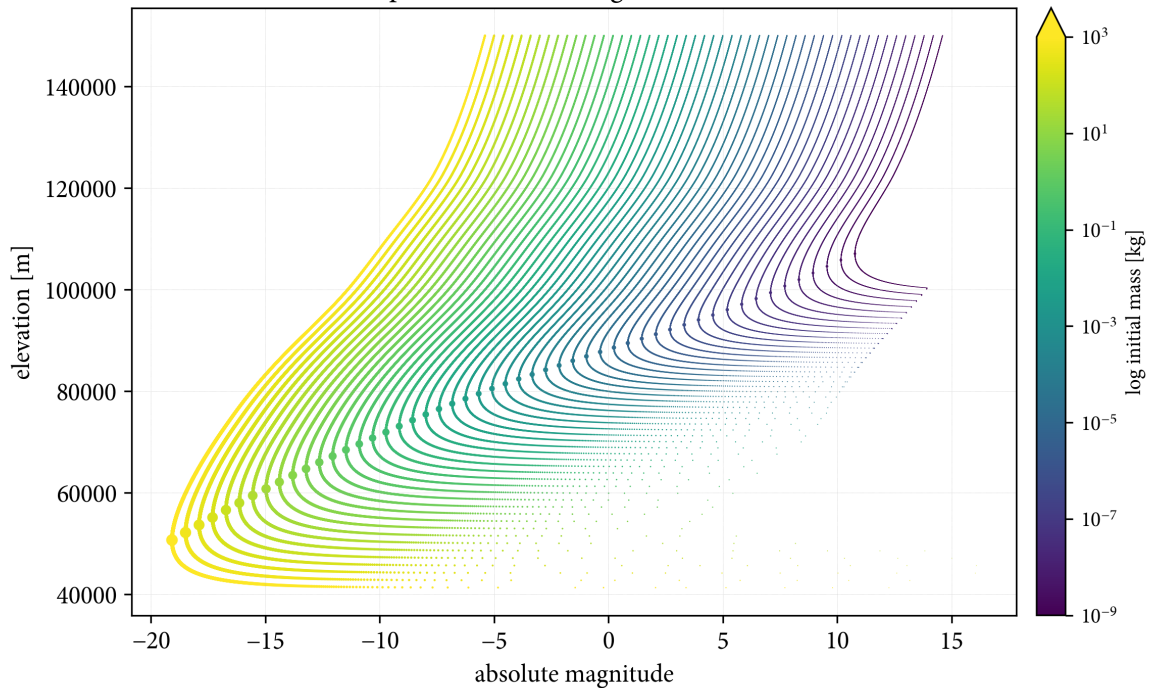
Parameters

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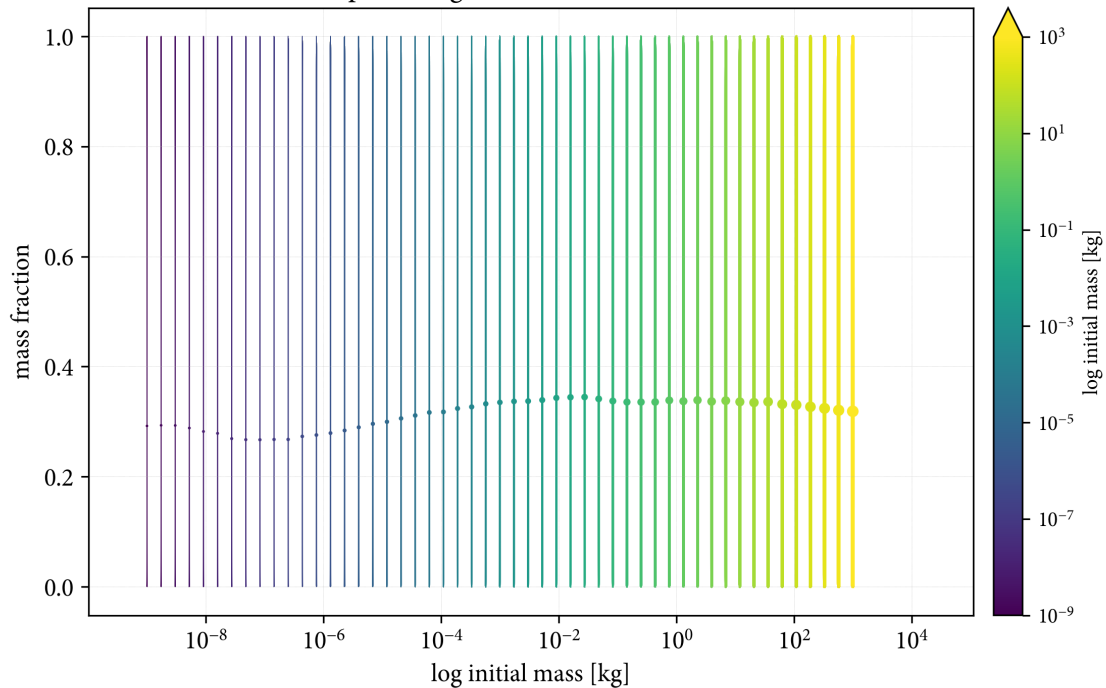
Parameters

- ▶ Perseids
 - ▶ material, velocity, ...
- ▶ entering atmosphere at 45°
- ▶ m_∞ from 1 μg to 1000 kg
- ▶ log spacing

North pole – absolute magnitude \times elevation



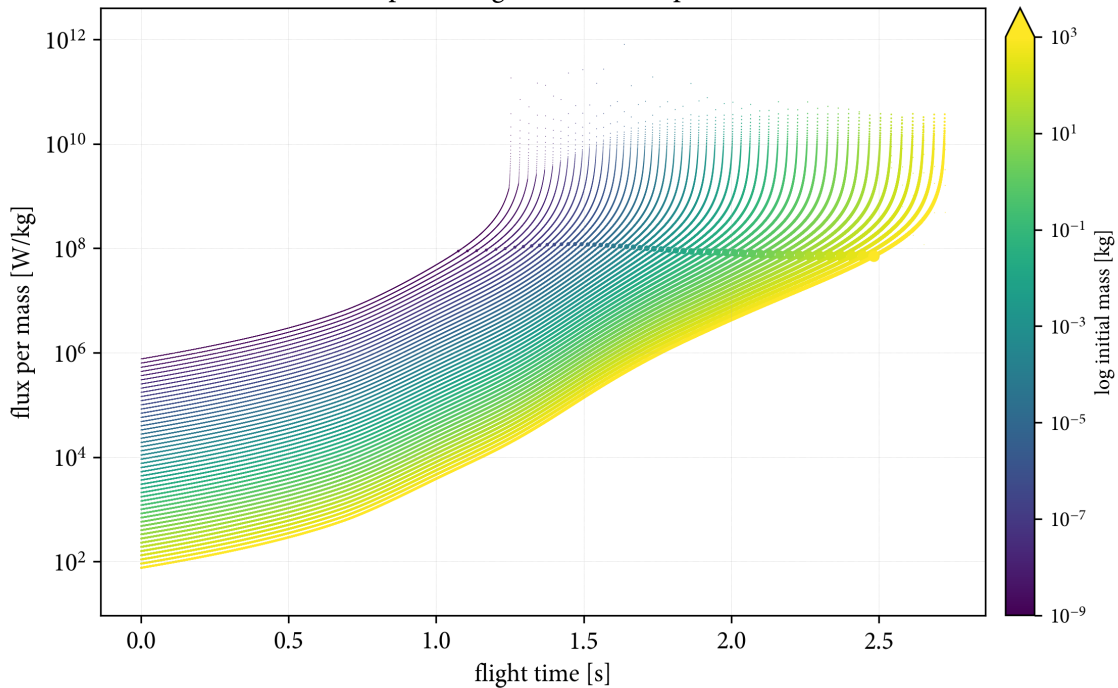
North pole – $\log \text{ initial mass} \times \text{mass fraction}$

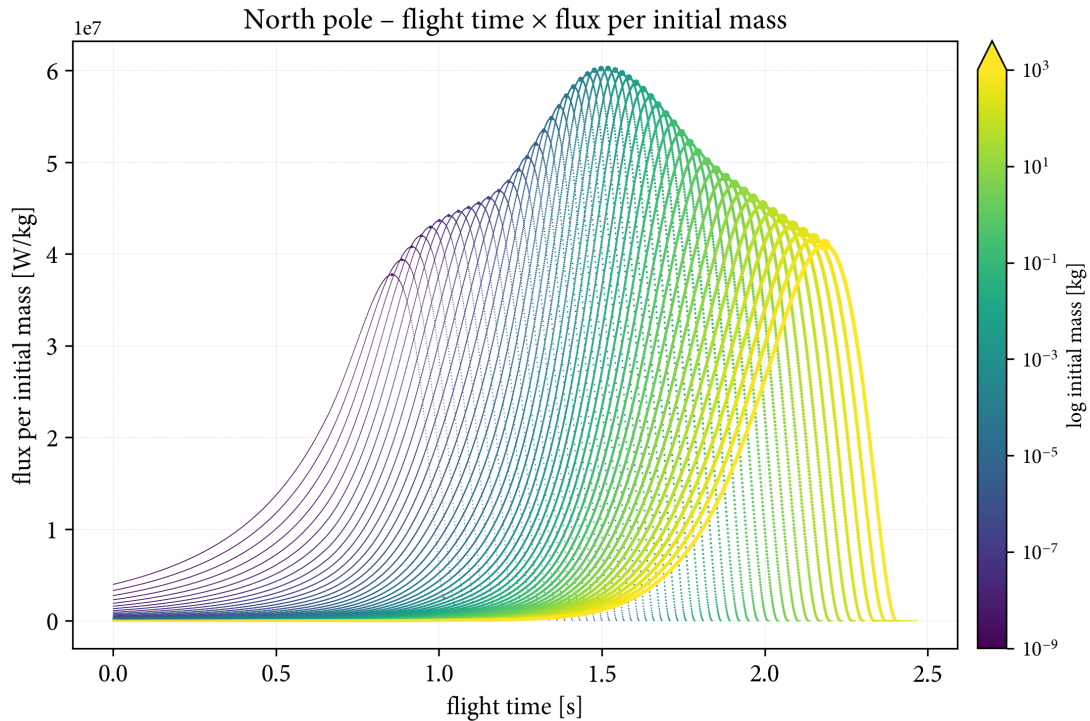


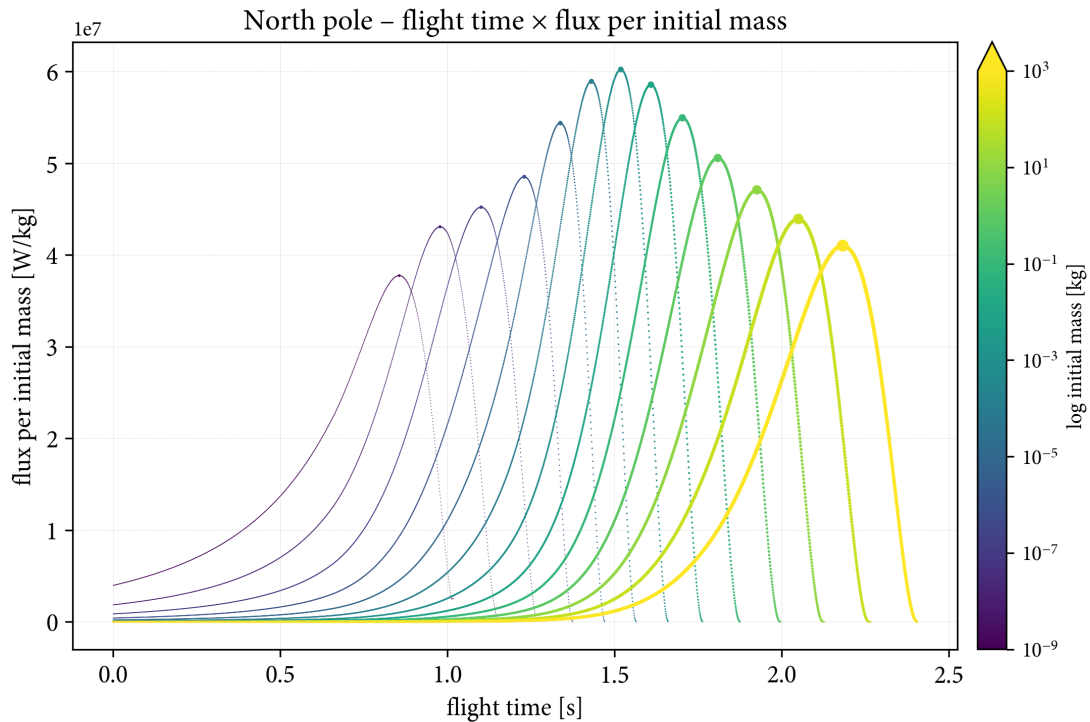
Mass fraction at maxlight

- ▶ maximum brightness at $1/3$ initial mass
- ▶ non-linear behaviour

North pole – flight time \times flux per mass







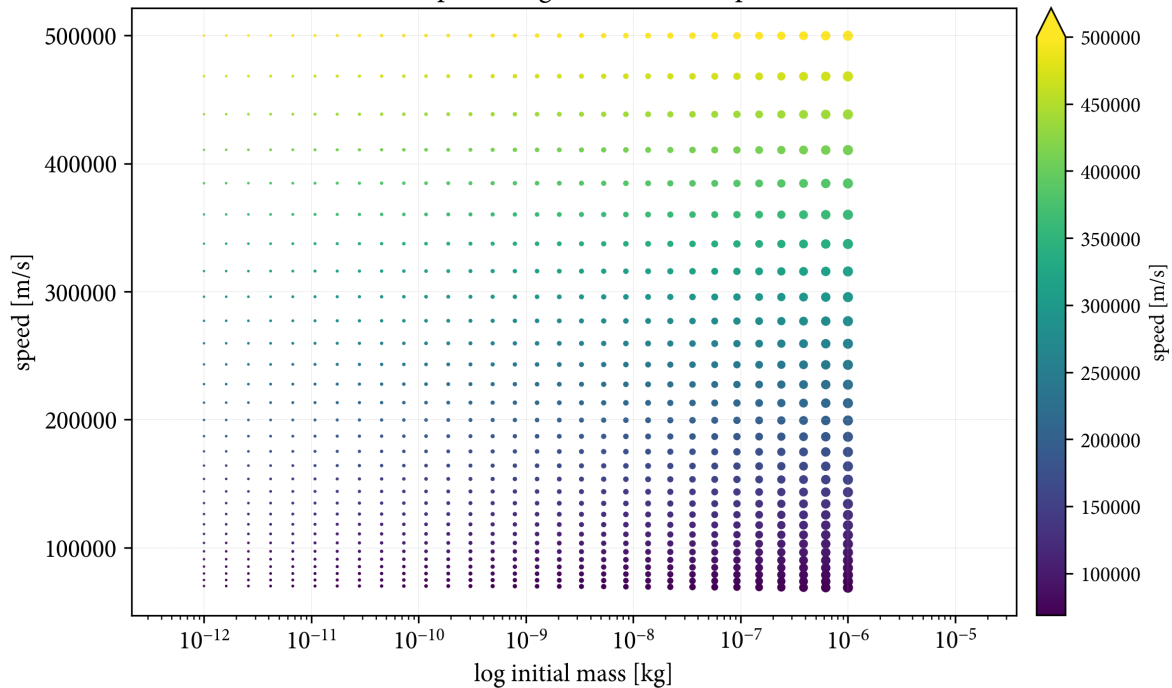
Hypervelocity

Hypervelocity

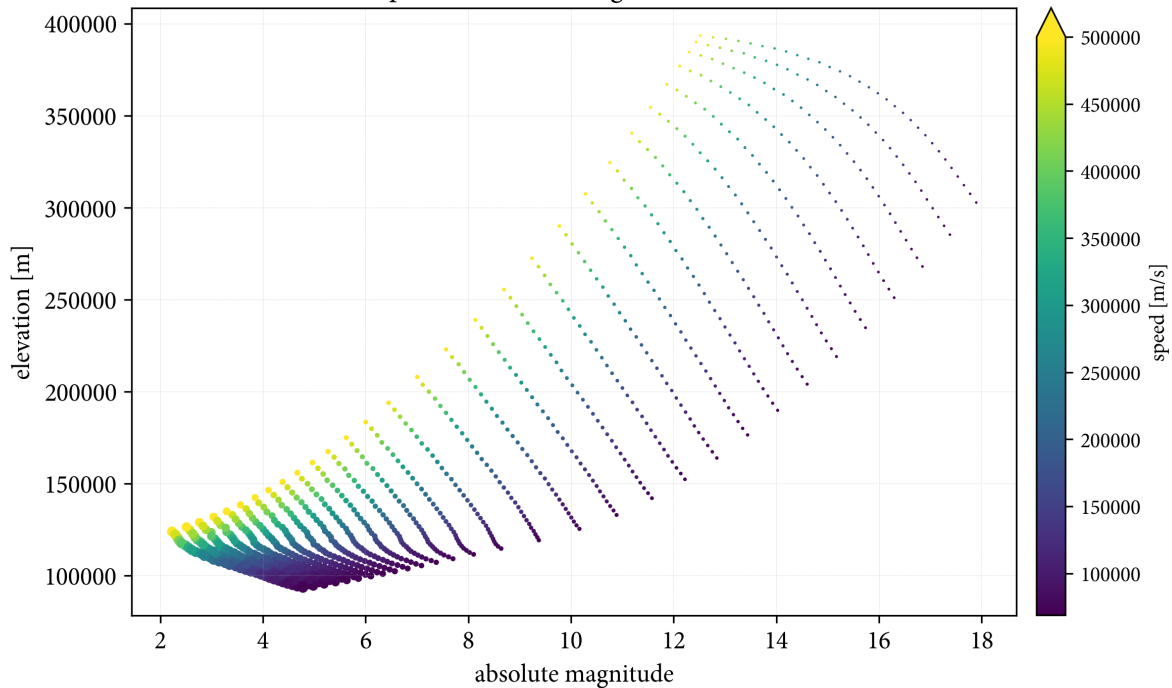
Revise **High geocentric velocity meteor ablation** (Hill et al., 2005)

- ▶ fluffy cometary material
- ▶ **speed** between $70 - 500 \text{ km s}^{-1}$
- ▶ **mass** between $1 \cdot 10^{-12} - 1 \cdot 10^{-6} \text{ kg}$

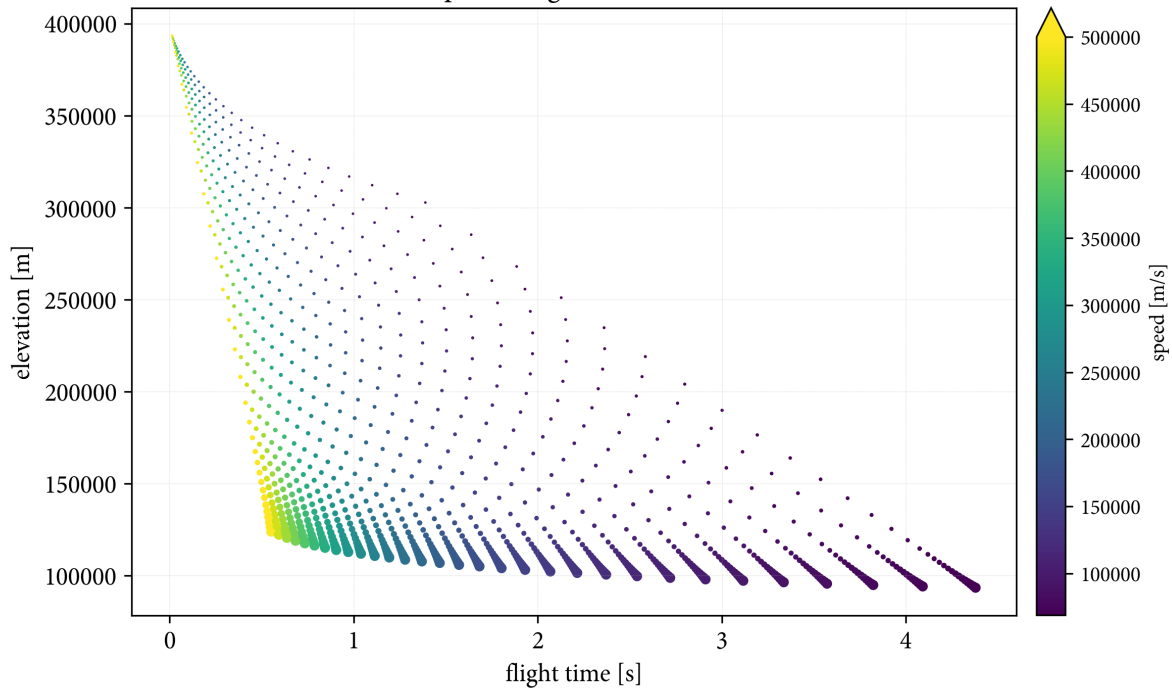
North pole – log initial mass \times speed



North pole – absolute magnitude \times elevation



North pole – flight time \times elevation



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- ▶ upper limit on flux
 - ▶ needs precise understanding of selection bias
- ▶ observation difficult
 - ▶ all-sky cameras
 - ▶ telescopes

Conclusion

Validation

- ▶ try to simulate real **individual meteors**
- ▶ revise the physics in the model
 - ▶ does not take body heating into account
 - ▶ very simple light emission model
 - ▶ not sensitive to Γ

Ultimate goal

- ▶ hopefully design a **better model** of meteor flight
 - ▶ compare with real meteors
 - ▶ use in flux estimations

Find motivation

...all play and no work makes Jack a mere toy.

Harry and Lucy Concluded
Maria Edgeworth (1825)

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 - ▶ <https://github.com/sesquideus/asmodeus>
- ▶ available for anyone (Python required)
- ▶ suggestions or comments always welcome

References

- ▶ **Öpik, E. J.:** Physics of meteor flight in the atmosphere. Interscience Publishers, 1958.
- ▶ **Hill, K. A. – Rogers, L. A. – Hawkes, R. L.:** High geocentric velocity meteor ablation. Astronomy & Astrophysics 444, 615–624 (2005)
- ▶ **Jones, W. – Halliday, I.:** Effects of Excitation and Ionization in Meteor Trains. MNRAS vol. 321, 2001, pp417–423.